Package 'lorentz'

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Type Package Title The Lorentz Transform in Relativistic Physics Version 1.1-2 Suggests knitr,testthat,rmarkdown,covr Imports quadform,tensor,magic,magrittr Maintainer Robin K. S. Hankin <hankin.robin@gmail.com> **Depends** R (>= 4.1.0) **Description** The Lorentz transform in special relativity; also the gyrogroup structure of three-velocities. Performs active and passive transforms and has the ability to use units in which the speed of light is not unity. Includes some experimental functionality for celerity and rapidity. For general relativity, see the 'schwarzschild' package. License GPL-3 **Encoding** UTF-8 VignetteBuilder knitr URL https://github.com/RobinHankin/lorentz, https://robinhankin.github.io/lorentz/ BugReports https://github.com/RobinHankin/lorentz/issues RoxygenNote 7.2.3 NeedsCompilation no

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The Lorentz Transform in Relativistic Physics

Description

The Lorentz transform in special relativity; also the gyrogroup structure of three-velocities. Performs active and passive transforms and has the ability to use units in which the speed of light is not unity. Includes some experimental functionality for celerity and rapidity. For general relativity, see the 'schwarzschild' package.

Details

The DESCRIPTION file:

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Suggests:	knitr,testthat,rmarkdown,covr

lorentz-package

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Maintainer:	Robin K. S. Hankin <hankin.robin@gmail.com></hankin.robin@gmail.com>
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Index of help topics:

Extract.3vel Ops.3vel as.matrix.3vel boost c.3vel	Extract or replace parts of three-velocity Arithmetic Ops Group Methods for 3vel objects Coerce 3-vectors and 4-vectors to a matrix Lorentz transformations Combine vectors of three-velocities and four-velocities into a single vector
celerity	Celerity and rapidity
comm_fail	Failure of commutativity and associativity using visual plots
coordnames	Coordinate names for relativity
cosines	Direction cosines
fourmom	Four momentum
fourvel	Four velocities
galileo	Classical mechanics; Newtonian approximation; infinite speed of light
gam	Gamma correction
gyr	Gyr function
lorentz-package photon	The Lorentz Transform in Relativistic Physics Photons
print.3vel	Print methods for three-velocities and four-velocities
r3vel	Random relativistic velocities
reflect	Mirrors
seq.3vel	seq method for three velocities
sol	Speed of light and Minkowski metric
threevel	Three velocities
transform	The energy-momentum tensor

Author(s)

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References

- Ungar 2006. "Thomas precession: a kinematic effect...". *European Journal of Physics*, 27:L17-L20.
- https://www.youtube.com/watch?v=9Y9CxiukURw&index=68&list=PL9_n3Tqzq9iWtgD8P0JFdnVUCZ_ zw60iB

Examples

```
u <- as.3vel(c(0.3,0.6,-0.1)) # u is a three-velocity</pre>
gam(u)
                               # relativistic gamma term for u
U <- as.4vel(u)
                               # U is a four-velocity
B1 <- boost(u)
                               # B1 is the Lorentz transform matrix for u
                              # Lorentz transform of zero 4-velocity (=-u)
B1 %*% c(1,0,0,0)
B2 <- boost(as.3vel(c(-0.1,0.8,0.3)))</pre>
B3 <- boost(as.3vel(c(-0.1,0.1,0.9))) # more boosts
Bi <- B1 %*% B2 # Bi is the boost for successive Lorentz transforms
pureboost(Bi)
                   # Decompose Bi into a pure boost...
                   # and an orthogonal matrix
orthog(Bi)
Bj <- B2 %*% B1
                  # B1 and B2 do not commute...
(B1 %*% B2) %*% B3
B1 %*% (B2 %*% B3)
                      # ...but composition *is* associative
## Three velocities and the gyrogroup
## Create some random three-velocities:
u <- r3vel(10)
v <- r3vel(10)</pre>
w <- r3vel(10)</pre>
u+v
           # Three-velocity addition is not commutative...
v+u
u+(v+w)
         # ... nor associative
(u+v)+w
```

boost

Description

Coerce 3-vectors and 4-vectors to a matrix. A convenience wrapper for unclass()

Usage

S3 method for class '3vel'
as.matrix(x, ...)
S3 method for class '4vel'
as.matrix(x, ...)

Arguments

Х	Object of class 3vel or 4vel
	Further arguments (currently ignored)

Author(s)

Robin K. S. Hankin

Examples

as.matrix(r3vel(5))
as.matrix(r4vel(5))

boost

Lorentz transformations

Description

Lorentz transformations: boosts and rotations

Usage

```
boost(u=0)
rot(u,v,space=TRUE)
is.consistent.boost(L, give=FALSE, TOL=1e-10)
is.consistent.boost.galilean(L, give=FALSE, TOL=1e-10)
pureboost(L,include_sol=TRUE)
orthog(L)
pureboost.galilean(L, tidy=TRUE)
orthog.galilean(L)
```

Arguments

u, v	Three-velocities, coerced to class 3vel. In function boost(), if u takes the special default value 0, this is interpreted as zero three velocity
L	Lorentz transform expressed as a 4×4 matrix
TOL	Numerical tolerance
give	Boolean with TRUE meaning to return the transformed metric tensor (which should be the flat-space $eta()$; qv) and default FALSE meaning to return whether the matrix is a consistent boost or not
space	Boolean, with default TRUE meaning to return just the spatial component of the rotation matrix and FALSE meaning to return the full 4×4 matrix transformation
tidy	In pureboost.galilean(), Boolean with default TRUE meaning to return a "tidy" boost matrix with spatial components forced to be a 3×3 identity matrix
include_sol	In function pureboost(), Boolean with default TRUE meaning to correctly account for the speed of light, and FALSE meaning to assume $c = 1$. See details

Details

Arguments u, v are coerced to three-velocities.

A rotation-free Lorentz transformation is known as a *boost* (sometimes a *pure boost*), here expressed in matrix form. Pure boost matrices are symmetric if c = 1. Function boost(u) returns a 4×4 matrix giving the Lorentz transform of an arbitrary three-velocity u.

Boosts can be successively applied with regular matrix multiplication. However, composing two successive pure boosts does not in general return a pure boost matrix: the product is not symmetric in general. Also note that boost matrices do not commute. The resulting matrix product represents a *Lorentz transform*.

It is possible to decompose a Lorentz transform L into a pure boost and a spatial rotation. Thus L = OP where O is an orthogonal matrix and P a pure boost matrix; these are returned by functions orthog() and pureboost() respectively. If the speed of light is not equal to 1, the functions still work but can be confusing.

Functions pureboost.galilean() and orthog.galilean() are the Newtonian equivalents of pureboost() and orthog(), intended to be used when the speed of light is infinite (which causes problems for the relativistic functions).

As noted above, the composition of two pure Lorentz boosts is not necessarily pure. If we have two successive boosts corresponding to u and v, then the composed boost may be decomposed into a pure boost of boost(u+v) and a rotation of rot(u,v).

The reason argument include_sol exists is that function orthog() needs to call pureboost() in an environment where we pretend that c = 1.

Value

Function boost() returns a 4×4 matrix; function rot() returns an orthogonal matrix.

Note

Function rot() uses crossprod() for efficiency reasons but is algebraically equivalent to boost(-u-v) %*% boost(u) %*% boost(v).

boost

Author(s)

Robin K. S. Hankin

References

- Ungar 2006. "Thomas precession: a kinematic effect...". European Journal of Physics, 27:L17-L20
- Sbitneva 2001. "Nonassociative geometry of special relativity". International Journal of Theoretical Physics, volume 40, number 1, pages 359–362
- Wikipedia contributors 2018. "Wigner rotation", Wikipedia, The Free Encyclopedia. https://en.wikipedia.org/w/index.php?title=Wigner_rotation&oldid=838661305. Online; accessed 23 August 2018

Examples

```
boost(as.3vel(c(0.4,-0.2,0.1)))
u <- r3vel(1)
v <- r3vel(1)</pre>
w <- r3vel(1)</pre>
boost(u) - solve(boost(-u)) # should be zero
boost(u) %*% boost(v) # not a pure boost (not symmetrical)
boost(u+v) # not the same!
boost(v+u) # also not the same!
u+v # returns a three-velocity
boost(u) %*% boost(v) %*% boost(w) # associative, no brackets needed
boost(u+(v+w)) # not the same!
boost((u+v)+w) # also not the same!
rot(u,v)
rot(v,u)
            # transpose (=inverse) of rot(u,v)
rot(u,v,FALSE) %*% boost(v) %*% boost(u)
boost(u+v)
            # should be the same.
orthog(boost(u) %*% boost(v)) - rot(u,v,FALSE) # zero to numerical precision
pureboost(boost(v) %*% boost(u)) - boost(u+v) # ditto
## Define a random-ish Lorentz transform
L <- boost(r3vel(1)) %*% boost(r3vel(1)) %*% boost(r3vel(1))
## check it:
```

```
## Not run:
              # needs emulator package
quad.form(eta(),L) # should be eta()
## End(Not run)
## More concisely:
                           # should be TRUE
is.consistent.boost(L)
## Decompose L into a rotation and a pure boost:
U <- orthog(L)
P <- pureboost(L)</pre>
L - U %*% P
                       # should be zero (L = UP)
crossprod(U)
                           # should be identity (U is orthogonal)
                           # should be zero (P is symmetric)
P - t(P)
## First row of P should be a consistent 4-velocity:
is.consistent.4vel(P[1,,drop=FALSE],give=TRUE)
```

c.3vel	Combine vectors of three-velocities and four-velocities into a single
	vector

Description

Combines its arguments recursively to form a vector of three velocities or four velocities

Usage

```
## S3 method for class '3vel'
c(...)
## S3 method for class '3cel'
c(...)
## S3 method for class '4vel'
c(...)
```

Arguments

... Vectors of three-velocities

Details

Returns a vector of three-velocities or four-velocities. These are stored as three- or four- column matrices; each row is a velocity.

Names are inherited from the behaviour of cbind(), not c().

celerity

Note

This function is used extensively in inst/distributive_search.R. For "c" as in celerity or speed of light, see sol().

Author(s)

Robin K. S. Hankin

See Also

sol

Examples

c(r3vel(3),r3vel(6,0.99))

celerity

Celerity and rapidity

Description

The celerity and rapidity of an object (experimental functionality)

Usage

```
## S3 method for class '3vel'
celerity(u)
## S3 method for class '4vel'
celerity_ur(d)
## S3 method for class '3vel'
rapidity(u)
## S3 method for class '4vel'
rapidity(u)
rapidity_ur(d)
as.3cel(x)
cel_to_vel(x)
vel_to_cel(x)
```

Arguments

u, x	Speed: either a vector of speeds or a vector of three-velocities or four-velocities
d	In functions celerity_ur() and rapidity_ur(), deficit of speed; speed of
	light minus speed of object

Details

The celerity corresponding to speed u is defined as $u\gamma$ and the rapidity is $c \cdot \operatorname{atanh}(u/c)$.

Functions celerity_ur() and rapidity_ur() are used for the ultrarelativistic case where speeds are very close to the speed of light. Its argument d is the deficit, that is, d = c - v where v is the speed of the transformation. Algebraically, celerity_ur(c-v) == celerity(v), but if d = 1 - v/c is small the result of celerity_ur() is more accurate than that of celerity().

Things get a bit sticky for celerity and rapidity if $c \neq 1$. The guiding principle in the package is to give the celerity and rapidity the same units as c, so if $u \ll c$ we have that all three of celerity(u), rapidity(u) and u are approximately equal. Note carefully that, in contrast, γ is dimensionless. Also observe that d in functions celerity_ur() and rapidity_ur() has the same units as c.

Author(s)

Robin K. S. Hankin

See Also

gam

Examples

```
u <- 0.1 # c=1
c(u,celerity(u),rapidity(u))
```

```
omgp <- 4.9e-24 # speed deficit of the Oh-My-God particle
c(celerity_ur(omgp),rapidity_ur(omgp))
```

```
snail <- 0.00275
c(snail,celerity(snail),rapidity(snail))</pre>
```

```
omgp <- omgp*sol()
c(celerity_ur(omgp),rapidity_ur(omgp))</pre>
```

sol(1)

comm_fail

Description

Relativistic addition of three-velocities is neither commutative nor associative, and the functions documented here show this visually.

Usage

```
comm_fail1(u, v, bold=5, r=1)
comm_fail2(u, v, bold=5, r=1)
ass_fail(u, v, w, bold=5, r=1)
my_seg(u,start=as.3vel(0), bold=5, ...)
```

Arguments

u,v,w,start	Three velocities. Arguments u and w are single-element three velocities, argument v is a vector. See the examples
bold	Integer specifying which vector element to be drawn in bold
r	Radius of dotted green circle, defaulting to 1 (corresponding to $c=1\).$ Use NA to suppress plotting of circle
	Further arguments, passed to arrows()

Value

These functions are called for their side-effect of plotting a diagram.

Note

The vignette lorentz gives more details and interpretation of the diagrams. Function my_seg() is an internal helper function.

Author(s)

Robin K. S. Hankin

Examples

```
u <- as.3vel(c(0.4,0,0))
v <- seq(as.3vel(c(0.4,-0.2,0)), as.3vel(c(-0.3,0.9,0)),len=20)
w <- as.3vel(c(0.8,-0.4,0))
comm_fail1(u=u, v=v)
comm_fail2(u=u, v=v)
ass_fail(u=u, v=v, w=w, bold=10)
```

coordnames

Description

Trivial function to set coordinate names to "t", "x", "y", "z".

Usage

```
coordnames(...)
flob(x)
```

Arguments

	Further arguments, currently ignored
x	A matrix

Details

Function coordnames() simply returns the character string c("t", "x", "y", "z"). It may be overwritten. Function flob() sets the row and columnnames of a 4×4 matrix to coordnames().

Note

If anyone can think of a better name than flob() let me know.

Author(s)

Robin K. S. Hankin

Examples

coordnames()

flob(diag(3))
flob(matrix(1,4,4))

You can change the names if you wish: coordnames <- function(x){letters[1:4]} flob(outer(1:4,1:4)) cosines

Description

Given a vector of three-velocities, returns their direction cosines

Usage

cosines(u, drop = TRUE)

Arguments

u	A vector of three-velocities
drop	Boolean, with default TRUE meaning to coerce return value from a one-row ma-
	trix to a vector, and FALSE meaning to consistently return a matrix

Author(s)

Robin K. S. Hankin

Examples

```
cosines(r3vel(7))
```

cosines(r3vel(1),drop=TRUE)
cosines(r3vel(1),drop=FALSE)

Extract.3vel

Extract or replace parts of three-velocity

Description

Extract or replace subsets of three-velocities

Arguments

х	A three-vector
index	elements to extract or replace
value	replacement value

Details

These methods (should) work as expected: an object of class 3vel is a three-column matrix with rows corresponding to three-velocities; a single argument is interpreted as a row number. Salient use-cases are u[1:5] <- u[1] and u[1] <- 0.

To extract a single component, pass a second index: u[,1] returns the x- component of the three-velocity.

Extraction functions take a drop argument, except for x[i] which returns a vec object.

Currently, u[] returns u but I am not sure this is desirable. Maybe it should return unclass(u) or perhaps c(unclass(u)).

Use idiom u[] <- x to replace entries of u elementwise.

Examples

```
u <- r3vel(10)
u[1:4]
u[5:6] <- 0
u[7:8] <- u[1]
u[,1] <- 0.1</pre>
```

fourmom

Four momentum

Description

Create and test for four-momentum

Usage

```
## S3 method for class '4mom'
Ops(e1, e2)
## S3 method for class '4mom'
sum(..., na.rm=FALSE)
vel_to_4mom(U,m=1)
p_to_4mom(p,E=1)
as.4mom(x)
is.4mom(x)
fourmom_mult(P,n)
fourmom_add(e1,e2)
```

fourmom

Arguments

x, P, e1, e2	Four-momentum
р	Three-momentum
E	Scalar; energy
U	Object coerced to four-velocity
m	Scalar; rest mass
n	Multiplying factor
,na.rm	Arguments sent to sum()

Details

Four-momentum is a relativistic generalization of three-momentum, with the object's energy as the first element. It can be defined as mU, where m is the rest mass and U the four-velocity. Equivalently, one can define four-momentum as $(E/c, p_x, p_y, p_z)$ where E is the energy and (p_x, p_y, p_z) the three-momentum.

Function vel_to_4mom() converts three-velocity to four-momentum, and function p_to_4mom()) converts a three-momentum to a four-momentum.

The function Ops.4mom() passes unary and binary arithmetic operators "+", "-" and "*" to the appropriate specialist function.

The package is designed so that natural R idiom may be used for physically meaningful operations such as combining momenta of different objects, using the conservation of four-momentum.

For the four-momentum of a photon, use as.photon().

Author(s)

Robin K. S. Hankin

See Also

boost,as.photon

Examples

```
# Define 5 random three velocities:
v <- r3vel(5)</pre>
```

convert to four-velocity: as.4vel(v)

```
# Now convert 'v' to four-momentum, specifying rest mass:
vel_to_4mom(v)  # 4mom of five objects with 3vel v, all unit mass
vel_to_4mom(v, 1:5)  # 4mom of five objects with 3vel v, masses 1-5
vel_to_4mom(v[1],1:5)  # 4mom of five objects with same 3vel, masses 1..5
# Now convert 'v' to four-momentum, specifying energy E:
p_to_4mom(v,E=1)
p_to_4mom(v,E=10)  # slower
```

fourvel

```
p_to_4mom(v,E=100) # even slower
# Four-momentum of objects moving closely parallel to the x-axis:
P <- vel_to_4mom(as.3vel(c(0.8,0,0)) + r3vel(7,0.01))
reflect(P)
reflect(P,c(1,1,1))
sum(P)
```

fourvel

Four velocities

Description

Create and test for four-velocities.

Usage

```
as.4vel(u)
is.consistent.4vel(U, give=FALSE, TOL=1e-10)
inner4(U,V=U)
to3(U)
```

Arguments

u	A vector of three-velocities
U, V	A vector of four-velocities
give	In function is.consistent.4vel(), Boolean with TRUE meaning to return $U \cdot U + c^2$, which is zero for a four-velocity, and default FALSE meaning to return whether the four-velocity is consistent to numerical precision
TOL	Small positive value used for tolerance

Details

Function as.4vel() takes a three-velocity and returns a four-velocity.

Given a four-vector V, function inner4() returns the Lorentz invariant $V^i V_i = \eta_{ij} V^i V^j$. This quantity is unchanged under Lorentz transforms. Note that function inner4() works for any four-vector, not just four-velocities. It will work for (eg) a four-displacement, a four-momentum vector or a four-frequency. In electromagnetism, we could have a four-current or a four-potential. If U is a four-velocity, then $U^i U_i = -c^2$; if U is a 4-displacement, then $U^i U_i$ is the squared interval. If P is the four-momentum of a photon then $P^i P_i = 0$.

Function to3() is a low-level helper function used when as . 3vel() is given a four-velocity.

Function is.consistent.4vel() checks for four-velocities being consistent in the sense that $U^i U_i = -c^2$. Giving this function a vector, for example, is.consistent.4vel(1:5), will return an error.

Compare the functions documented here with boost(), which returns a 4×4 transformation matrix (which also includes rotation information).

galileo

Author(s)

Robin K. S. Hankin

See Also

boost

Examples

```
a <- r3vel(10)
as.4vel(a)  # a four-velocity
as.3vel(as.4vel(a))-a  # zero to numerical precision
inner4(as.4vel(a))  # -1 to numerical precision
stopifnot(all(is.consistent.4vel(as.4vel(a))))
## check Lorentz invariance of dot product:
U <- as.4vel(r3vel(10))</pre>
```

```
V <- as.4vel(r3vel(10))
B <- boost(as.3vel(1:3/10))
frame1dotprod <- inner4(U, V)
frame2dotprod <- inner4(U %*% B, V %*% B)
max(abs(frame1dotprod-frame2dotprod)) # zero to numerical precision</pre>
```

galileo

Classical mechanics; Newtonian approximation; infinite speed of light

Description

The Lorentz transforms reduce to their classical limit, the Galilean transforms, if speeds are low compared with c. Package idiom for working in a classical framework is to use an infinite speed of light: sol(Inf). Here I show examples of this.

Author(s)

Robin K. S. Hankin

See Also

boost

Examples

```
sol(Inf)
boost(as.3vel(1:3))
as.3vel(1:3) + as.3vel(c(-1,4,5))  # classical velocity addition
rot(as.3vel(1:3),as.3vel(c(-4,5,2)))  # identity matrix
B <- boost(as.3vel(1:3))
orthog(B) %*% pureboost(B)  # should be B
sol(1)
```

gam

Gamma correction

Description

Lorentz gamma correction term in special relativity

Usage

```
## S3 method for class '3vel'
speed(u)
## S3 method for class '4vel'
speed(u)
speedsquared(u)
gam(u)
gamm1(u)
## S3 method for class '3vel'
gam(u)
## S3 method for class '3cel'
gam(u)
## S3 method for class '4vel'
gam(u)
## S3 method for class '3vel'
gamm1(u)
## S3 method for class '4vel'
gamm1(u)
gam_ur(d)
```

Arguments

u	Speed: either a vector of speeds or a vector of three-velocities or four-velocities
d	In function gam_ur(), deficit of speed; speed of light minus speed of object

Details

Function speed(u) returns the speed of a 3vel object or 4vel object.

Function gam(u) returns the Lorentz factor

$$\frac{1}{\sqrt{1-\mathbf{u}\cdot\mathbf{u}/c^2}}$$

Function gamm1(u) returns the Lorentz factor minus 1, useful for slow speeds when larger accuracy is needed (much like expm1()); to see the R idiom, type "gamm1.3vel" at the commandline. Function gamm1() is intended to work with 3vel objects or speeds. The function will take a 4-velocity, but this is not recommended as accuracy is lost (all it does is return the time component of the 4-velocity minus 1).

Function gam_ur() is used for the ultrarelativistic case where speeds are very close to the speed of light (the function is named for "gamma, ultrarelativistic"). Its argument d is the deficit, that is, c - v where v is the speed of the transformation. Algebraically, gam_ur(c-v) == gam(v), but if d is small compared to c the result is more accurate.

Function speedsquared(u) returns the square of the speed of a 3vel object. Use this to avoid taking a needless square root.

Author(s)

Robin K. S. Hankin

Examples

```
gam(seq(from=0,by=0.1,len=10))
gam(r3vel(6,0.7))
```

```
x <- as.3vel(c(0.1,0.4,0.5))
speed(x)</pre>
```

```
gam(speed(x)) # works, but slow and inaccurate
gam(x) # recommended: avoids needless coercion
```

```
## Use SI units and deal with terrestrial speeds. Use gamm1() for this.
sol(299792458)
sound <- 343 # speed of sound in SI
gam(sound)
gam(sound)-1
gamm1(sound) # gamm1() gives much higher precision
```

```
## For the ultrarelativistic case of speeds very close to the speed of
## light, use gam_ur():
```

```
20
```

```
sol(1)  # revert to relativistic units
gam(0.99) - gam_ur(0.01) # zero to numerical accuracy
omgp <- 4.9e-24 # speed deficit of the Oh-My-God particle
gam(1-omgp) # numeric overflow
gam_ur(omgp) # large but finite</pre>
```

gyr

Gyr function

Description

Relativistic addition of three velocities

Usage

gyr(u, v, x)
gyr.a(u, v, x)
gyrfun(u, v)

Arguments

u, v, x Three-velocities, objects of class 3vel

Details

Function gyr(u,v,x) returns the three-vector gyr[u, v]x.

Function gyrfun(u, v) returns a function that returns a three-vector; see examples.

The speed of light (1 by default) is not used directly by these functions; set it with sol().

Note

Function gyr() is slightly faster than gyr.a(), which is included for pedagogical reasons. Function gyr() is simply add3(neg3(add3(u,v)),add3(u,add3(v,x))) while function gyr.a() uses the slower but more transparent idiom -(u+v) + (u+(v+x))

Author(s)

Robin K. S. Hankin

Ops.3vel

References

- Ungar 2006. "Thomas precession: a kinematic effect of the algebra of Einstein's velocity addition law. Comments on 'Deriving relativistic momentum and energy: I. Three-dimensional case". European Journal of Physics, 27:L17-L20.
- Sbitneva 2001. "Nonassociative geometry of special relativity". International Journal of Theoretical Physics, volume 40, number 1, pages 359–362

Examples

```
u <- r3vel(10)
v <- r3vel(10)</pre>
w <- r3vel(10)</pre>
x <- as.3vel(c(0.4,0.1,-0.5))</pre>
y <- as.3vel(c(0.1,0.2,-0.7))</pre>
z <- as.3vel(c(0.2,0.3,-0.1))</pre>
gyr(u,v,x) # gyr[u,v]x
f <- gyrfun(u,v)</pre>
g <- gyrfun(v,u)
f(x)
f(r3vel(10))
f(g(x)) - x
                         # zero, by eqn 9
g(f(x)) - x
                       # zero, by eqn 9
(x+y) - f(y+x)
                       # zero by eqn 10
(u+(v+w)) - ((u+v)+f(w)) # zero by eqn 11
# Following taken from Sbitneva 2001:
rbind(x+(y+(x+z)) ,
                       (x+(y+x))+z) # left Bol property
rbind((x+y)+(x+y) ,
                                       # left Bruck property
                       x+(y+(y+x)))
sol(299792458) # speed of light in SI
as.3vel(c(1000,3000,1000)) + as.3vel(c(1000,3000,1000))
## should be close to Galilean result
sol(1) # revert to default c=1
```

Ops.3vel

Description

Arithmetic operations for three-velocities

Usage

```
## S3 method for class '3vel'
Ops(e1, e2)
## S3 method for class '4vel'
Ops(e1, e2)
massage3(u,v)
neg3(u)
prod3(u,v=u)
add3(u,v)
dot3(v,r)
```

Arguments

e1, e2, u, v	Objects of class "3vel", three-velocities
r	Scalar value for circle-dot multiplication

Details

The function Ops.3vel() passes unary and binary arithmetic operators "+", "-" and "*" to the appropriate specialist function.

The most interesting operators are "+" and "*", which are passed to add3() and dot3() respectively. These are defined, following Ungar, as:

$$\mathbf{u} + \mathbf{v} = \frac{1}{1 + \mathbf{u} \cdot \mathbf{b}/c^2} \left\{ \mathbf{u} + \frac{1}{\gamma_{\mathbf{u}}} \mathbf{v} + \frac{1}{c^2} \frac{\gamma_{\mathbf{u}}}{1 + \gamma_{\mathbf{u}}} \left(\mathbf{u} \cdot \mathbf{v} \right) \mathbf{u} \right\}$$

and

$$r \odot \mathbf{v} = c \tanh\left(r \tanh^{-1} \frac{||\mathbf{v}||}{c}\right) \frac{\mathbf{v}}{||\mathbf{v}||}$$

where \mathbf{u} and \mathbf{v} are three-vectors and r a scalar. Function dot3() has special dispensation for zero velocity and does not treat NA entries entirely consistently.

Arithmetic operations, executed via Ops.4vel(), are not defined on four-velocities.

The package is designed so that natural R idiom may be used for three velocity addition, see the examples section.

Value

Returns an object of class 3vel, except for prod3() which returns a numeric vector.

photon

Examples

```
u <- as.3vel(c(-0.7, 0.1,-0.1))
v <- as.3vel(c( 0.1, 0.2, 0.3))</pre>
w <- as.3vel(c( 0.5, 0.2,-0.3))</pre>
                 # random three velocities
x <- r3vel(10)
y <- r3vel(10) # random three velocities</pre>
u+v
      # add3(u,v)
      # add3(u,neg3(v))
u-v
      # neg3(v)
-v
gyr(u,v,w)
## package is vectorized:
u+x
x+y
f \leq gyrfun(u,v)
g <- gyrfun(v,u)</pre>
f(g(x)) - x
               # should be zero by eqn10
g(f(x)) - x
(u+v) - f(v+u)
                                   # zero by eqn 10
(u+(v+w)) - ((u+v)+f(w))
                                  # zero by eqn 11
((u+v)+w) - (u+(v+g(w)))
                                    # zero by eqn 11
## NB, R idiom is unambiguous. But always always ALWAYS use brackets.
## Ice report in lat 42.n to 41.25n Long 49w to long 50.30w saw much
## heavy pack ice and great number large icebergs also field
## ice. Weather good clear
## -u+v == (-u) + v == neg3(u) + v == add3(neg3(u),v)
## u+v+w == (u+v)+w == add3(add3(u,v),w)
```

```
photon
```

Photons

Description

Various functionality to deal with the 4-momentum of a photon

photon

Usage

is.consistent.nullvec(N,TOL=1e-10)
as.photon(x,E=1)

Arguments

Ν	Four-momentum to be tested for nullness
TOL	tolerance
x	Vector of three-velocities
E	Energy, a scalar

Details

Returns the four-momentum of a photon.

Author(s)

Robin K. S. Hankin

See Also

4mom,reflect

Examples

```
## A bunch of photons all approximately parallel to the x-axis:
as.photon(as.3vel(cbind(0.9,runif(10)/1000,runif(10)/1000)))
```

```
## mirror ball:
jj <- matrix(rnorm(30),10,3)
disco <- sweep(matrix(rnorm(30),10,3),1,sqrt(rowSums(jj^2)),`/`)
p <- as.photon(c(1,0,0))
reflect(p,disco)
table(reflect(p,disco)[,2]>0) # should be TRUE with probability sqrt(0.5)
## relativistic disco; mirror ball moves at 0.5c:
B <- boost(as.3vel(c(0.5,0,0)))</pre>
```

```
p |> tcrossprod(B) |> reflect(disco) |> tcrossprod(solve(B))
```

print.3vel

Description

Print methods for three-velocities

Usage

```
## S3 method for class '3vel'
print(x, ...)
## S3 method for class '3cel'
print(x, ...)
## S3 method for class '4vel'
print(x, ...)
## S3 method for class '4mom'
print(x, ...)
```

Arguments

Х	Vector of three-velocities
	Further arguments, currently ignored

Value

Returns a vector of three-velocities

Author(s)

Robin K. S. Hankin

Examples

r3vel(10)

r3vel

Random relativistic velocities

Description

Generates random three-velocities or four-velocities, optionally specifying a magnitude

Usage

r3vel(n=7, r = NA) r4vel(...) rboost(r = NA)

Arguments

n	Number of three- or four- velocities to generate
r	Absolute value of the three-velocities, with default NA meaning to sample uniformly from the unit ball
	Arguments passed to r3vel()

Details

Function r3vel() returns a random three-velocity. Function r4vel() is a convenience wrapper for as.4vel(r3vel()).

Function rboost() returns a random 4×4 Lorentz boost matrix, drawn from the connected component. If given r=0, then a transform corresponding to a random rotation will be returned.

Value

Returns a vector of three- or four- velocities.

Note

If the speed of light is infinite, these functions require a specified argument for r.

It is not entirely trivial to sample *uniformly* from the unit ball or unit sphere, but it is not hard either.

Author(s)

Robin K. S. Hankin

Examples

```
r3vel()
a <- r3vel(10000)
b <- r3vel(1000,0.8)
u <- as.3vel(c(0,0,0.9))
pairs(unclass(u+a),asp=1)
pairs(unclass(a+u),asp=1)
is.consistent.boost(rboost())
sol(299792458)  # switch to SI units
sound <- 343  # speed of sound in metres per second
r3vel(100,343)  # random 3-velocities with speed = 343 m/s
sol(1)  # return to default c=1</pre>
```

reflect

Description

Plane mirrors in special relativity

Usage

reflect(P,m,ref=1)

Arguments

Р	Vector of four-momenta
m	Orientation of mirror, expressed as a three-vector
ref	Coefficient of reflectivity of the mirror

Value

Takes a four-momentum and returns the four-momentum after reflection. Will handle objects or photons.

Note

All four-momenta are measured in the rest frame of the mirror, but it is easy to reflect from moving mirrors; see examples.

However, note that the ref argument is designed to work with photons only, where it is conceptually the percentage of photons reflected and not absorbed by the mirror. If ref is less than unity, odd results are given for four momenta of nonzero restmass objects.

Author(s)

Robin K. S. Hankin

See Also

photon

Examples

```
## We will reflect some photons from an oblique mirror moving at half
## the speed of light.
## First create 'A', a bunch of photons all moving roughly along the x-axis:
A <- as.photon(as.3vel(cbind(0.9,runif(10)/1000,runif(10)/1000)))
## Now create 'm', a mirror oriented perpendicular to c(1,1,1):
m <- c(1,1,1)</pre>
```

```
## Reflect the photons in the mirror:
reflect(A,m)
## Reflect the photons in a series of mirrors:
A |> reflect(m) |> reflect(1:3) |> reflect(3:1)
## To reflect from a moving mirror we need to transform to a frame in
## which the mirror is at rest, then transform back to the original
## frame. First create B, a boost representing the mirror's movement
## along the x-axis at speed c/2:
B <- boost(as.3vel(c(0.5,0,0)))
## Transform to the mirror's rest frame:
A %*% t(B)
## NB: in the above, take a transpose because the *rows* of A are 4-vectors.
## Then reflect the photons in the mirror:
reflect(A %*% t(B),m)
## Now transform back to the original rest frame (NB: active transform):
```

A |> tcrossprod(B) |> reflect(m) |> tcrossprod(solve(B))

seq.3vel

seq method for three velocities

Description

Simplified version of seq() for three-velocities.

Usage

```
## S3 method for class '3vel'
seq(from, to, len, ...)
```

Arguments

from, to	Start and end of sequence
len	Length of vector returned
	Further arguments (currently ignored)

Details

seq(a,b,n) returns a + t*(-b+a) where t is numeric vector seq(from=0, to=1, len=n).

This definition is one of several plausible alternatives, but has the nice property that the first and last elements are exactly equal to a and b respectively.

Author(s)

Robin K. S. Hankin

Examples

```
a <- as.3vel(c(4,5,6)/9)
b <- as.3vel(c(-5,6,8)/14)
x <- seq(a,b,len=9)
x[1]-a # should be zero
x[9]-b # should be zero
jj <- a + seq(0,1,len=9)*(b-a)
jj-x # decidedly non-zero
```

sol

Speed of light and Minkowski metric

Description

Getting and setting the speed of light

Usage

```
sol(c)
eta(downstairs=TRUE)
ptm(to_natural=TRUE, change_time=TRUE)
```

Arguments

с	Scalar, speed of light. If missing, return the speed of light
downstairs	Boolean, with default TRUE meaning to return the covariant metric tensor g_{ij} with two downstairs indices, and FALSE meaning to return the contravariant version g^{ij} with two upstairs indices
to_natural, change_time	
	Pooleon specifying the neture of the passive transform matrix

Boolean, specifying the nature of the passive transform matrix

30

In the context of an R package, the symbol "c" presents particular problems. In the **lorentz** package, the speed of light is denoted "sol", for 'speed of light'. You can set the speed of light with sol(x), and query it with sol(); see the examples. An infinite speed of light is sometimes useful for Galilean transforms.

The speed of light is a global variable, governed by options("c"). If NULL, define c=1. Setting showSOL to TRUE makes sol() change the prompt to display the speed of light which might be useful.

Function eta() returns the Minkowski flat-space metric

diag
$$(-c^2, 1, 1, 1)$$
.

Note that the top-left element of eta() is $-c^2$, not -1.

Function ptm() returns a passive transformation matrix that converts displacement vectors to natural units (to_natural=TRUE) or from natural units (to_natural=FALSE). Argument change_time specifies whether to change the unit of time (if TRUE) or the unit of length (if FALSE).

Note

Typing "sol(299792458)" is a lot easier than typing "options("c"=299792458)", which is why the package uses the idiom that it does.

In a R-devel discussion about options for printing, Martin Maechler makes the following observation: "Good programming style for functions according to my book is to have them depend only on their arguments, and if a global option really (really? think twice!) should influence behavior, there should be arguments of the function which have a default determined by the global option"

I think he is right in general, but offer the observation that the speed of light depends on the units chosen, and typically one fixes one's units once and for all, and does not subsequently change them. This would indicate (to me at least) that a global option would be appropriate. Further, there *is* a default, c = 1, which is returned by sol() if the option is unset. This is not just a "default", though: it is used in the overwhelming majority of cases. Indeed, pedagogically speaking, one learning objective from the package is that units in which $c \neq 1$ are difficult, awkward, and unnatural. In the package R code, the *only* place the speed of light option is accessed is via sol(). Similar arguments are presented in the **clifford** package at signature.Rd.

Looking again at Martin's observation he seems to be suggesting that something along the lines of

gam <- function(u, c=1){1/sqrt(1-u^2/c^2)}</pre>

But this is asking for trouble:

c(gam(0.4,c=1),gam(0.4,c=10))

which is meaningless at best and misleading at worst.

threevel

Author(s)

Robin K. S. Hankin

Examples

```
sol()  # returns current speed of light
sol(299792458)  # use SI units
sol()  # speed of light now SI value
eta()  # note [t,t] term
u <- as.3vel(c(100,200,300))  # fast terrestrial speed, but not relativistic
boost(u)  # boost matrix practically Galilean
is.consistent.boost(boost(u))  # should be TRUE
sol(1)  # revert to relativistic units
```

threevel

Three velocities

Description

Create and test for three-velocities, 3vel objects.

Usage

```
`3vel`(n)
threevel(n)
as.3vel(x)
is.3vel(x)
## S3 method for class 'vec'
length(x)
## S3 method for class 'vec'
names(x)
## S3 replacement method for class 'vec'
names(x) <- value</pre>
```

Arguments

n	In function 3vel(), number of three velocities to create
x, value	Vectors of three-velocities

Note

Class vel is a virtual class containing classes 3vel and 4vel. Function threevel() is a convenience wrapper for 3vel().

transform

Author(s)

Robin K. S. Hankin

Examples

```
U <- r4vel(7)
as.4vel(as.3vel(U)) # equal to U, to numerical precision
x <- as.3vel(1:3/4)
u <- as.3vel(matrix(runif(30)/10,ncol=3))
names(u) <- letters[1:10]
x+u
u+x # not equal</pre>
```

transform

The energy-momentum tensor

Description

Various functionality to deal with the stress-energy tensor in special relativity.

Usage

```
perfectfluid(rho,p,u=0)
dust(rho,u=0)
photongas(rho,u=0)
transform_dd(TT, B)
transform_ud(TT, B)
transform_uu(TT, B)
raise(TT)
lower(TT)
```

Arguments

TT	A second-rank tensor with indices either downstairs-downstairs, downstairs- upstairs, or upstairs-upstairs
В	A boost matrix
rho, p, u	Density, pressure, and four-velocity of the dust

transform

Details

Function perfectfluid() returns the stress-energy tensor, with two upstairs indices, for a perfect fluid with the conditions specified. No checking for physical reasonableness (eg the weak energy condition) is performed: caveat emptor!

Function dust() is a (trivial) function that returns the stress-energy tensor of a zero-pressure perfect fluid (that is, dust). Function photongas() returns the stress-energy tensor of a photon gas. They are here for discoverability reasons; both are special cases of a perfect fluid.

Functions transform_dd() et seq transform a second-rank tensor using the Lorentz transform. The letters "u" or "d" denote the indices of the tensor being upstairs (contravariant) or downstairs (covariant). The stress-energy tensor is usually written with two upstairs indices, so use transform_uu() to transform it.

Function lower() lowers both indices of a tensor with two upstairs indices. Function raise() raises two downstairs indices. These two functions have identical R idiom but do not return identical values if $c \neq 1$.

Author(s)

Robin K. S. Hankin

Examples

```
perfectfluid(10,1)
u <- as.3vel(c(0.4,0.4,0.2))
## In the following, LHS is stationary dust and RHS is dust moving at
## velocity 'u', but transformed to a frame also moving at velocity 'u':
LHS <- dust(1)
RHS <- transform_uu(dust(1,u),boost(u))</pre>
max(abs(LHS-RHS)) # should be small
## In the following, negative sign needed because active/passive
## difference:
LHS <- dust(1,u)
RHS <- transform_uu(dust(1),boost(-u))</pre>
max(abs(LHS-RHS)) # should be small
## Now test behaviour when c!=1:
sol(299792458)
perfectfluid(1.225,101325) # air at STP
LHS <- transform_uu(perfectfluid(1.225,101325),boost(as.3vel(c(1000,0,0))))
```

```
LHS <- transform_uu(perfectfluid(1.225,101325),boost(as.svel(c(1000,0,0))))
RHS <- perfectfluid(1.225,101325)
LHS-RHS # should be small</pre>
```

transform

```
sol(10)
u <- as.3vel(4:6)
LHS <- photongas(1,u)
RHS <- transform_uu(photongas(1),boost(-u))
LHS-RHS # should be small
```

```
B1 <- boost(r3vel(1)) %*% boost(r3vel(1))
B2 <- boost(r3vel(1)) %*% boost(r3vel(1))
LHS <- transform_uu(transform_uu(dust(1),B1),B2)
RHS <- transform_uu(dust(1),B2 %*% B1)  # note order
LHS-RHS  # should be small</pre>
```

remember to re-set c:
sol(1)

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